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# Dynamical Low-Rank Approximation for Plasma Kinetic Model Reduction

Jack Coughlin, Jingwei Hu

November 19, 2021

# Outline

Approximation for Plasma Kinetic Model Reduction

Dynamical Low-Rank

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#### Plasmas and Kinetic Theory

2 Moment methods

3 Dynamical Low-Rank Approximation

4 Asymptotic-preserving low-rank methods

#### **5** Conclusion

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### What are plasmas?

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Plasmas are ionized, diffuse gases composed of charged particles (electrons and ions), and subject to electromagnetic forces. They are hard to model, but even harder to understand without computational modeling.

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# Two applications of plasma modeling

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#### Nuclear fusion requires taming a bewildering array of plasma instabilities.

10.7. <u>Name index</u> In this index the number preceding the class symbol refers to the corresponding page in the Appendix.

Absolute instability: Accounts instabilities Accounts unwe instability in a partially ionited plasme Alfords unve (fire hose) instability

Amplitude dispersion instability of whistlers

Amplitude disponsion instabilities; relativistic modes Anisotropic temperature instability:

Ballooning isstability Boom-contrifugal instability

Boss-plasma instability:

Bernstein-Greene-Freskal wave instability Backling instability

Bulge instability: Bunching instability: Buncan instability:

Contrifugal instability: Correnkov instability Collective instability

Collective electrostatic instabilities in a two--dimensional field

Collisional drift instability: Collisionless gravitation Esstability:

Colligionless tearing

see two-stream instability
 see two-stream instability
 see fatter, Soyleigh-Toylor, and gravitation instabilities
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see velocity space instability

bility new gravitation instability; collisionless modes

ses tearing instability;

Figure: Source [9]

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#### Plasmas and Kinetic Theory

# Two applications of plasma modeling

Nuclear fusion requires taming a bewildering array of plasma instabilities

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Collective electrostatic instabilition in a two-dimensional field

Colligional drift insta-Cellisicaless gravitation instability:

Collisionless tearing instability:

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|------|--|
|      | see velocity space instability                               |
| 47.  | (FW1C1SoTEDE1ELME1A12)                                       |
| 12.  | (PETC'Ss'VTRE'E'S'DERPATE)                                   |
| 111. | see two-styces instability<br>(PV'C'S'ITB'E'L'M'tEp'A'I)     |
| 42.  | (PV*C*SeTEM*ELRE*A*1)  |
|      | see nounage instability                                      |
|      | see two-stream instability                                   |
|      | see two-stream instability                                   |
|      | nos flute, Rayleigh-Taylor,<br>and gravitation instabilities |
| 2.   | (PVC'Sa'v70b'f'TLMEpA'I)                                     |
|      | see Sector 3.4.4.  |

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nos drift-dissipativo insta

see gravitation instability; ses tearing instability; collisionless mode-

Figure: Source [9]

Space is filled with plasma and magnetic fields.



Figure: Europa's induction response to Jupiter's time-varying magnetic field provides strong evidence of a liquid ocean. [1]

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# The Boltzmann equation

$$\partial_t f_{\alpha} + \mathbf{v} \cdot \nabla_{\mathbf{x}} f_{\alpha} + F \cdot \nabla_{\mathbf{v}} f_{\alpha} = \left(\frac{\partial f_{\alpha}}{\partial t}\right)_{\mathsf{coll}} = \sum_{\beta} S(f_{\alpha}, f_{\beta}).$$

- Indexed by particle species (lpha,eta)
- $f_{\alpha}(x,v,t)$  a probability distribution
- $F = E + v \times B$  couples with Maxwell's equations
- Nonlinearity comes from field coupling and from  $S(f_{\alpha}, f_{\alpha})$ .

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Posed in six dimensions!

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### The collision operator

... deserves its own slide.

Many collision models are used, but all satisfy a few invariants. Important properties of  $S(f_{\alpha}, f_{\beta})$ 

• Conservation of mass, momentum, and energy:

$$\begin{bmatrix} \int f_{\alpha} dv_{\alpha} + \int f_{\beta} dv_{\beta}, \\ \int f_{\alpha} m_{\alpha} v_{\alpha} dv_{\alpha} + \int f_{\beta} m_{\beta} v_{\beta} dv_{\beta}, \\ \int f_{\alpha} m_{\alpha} \frac{|v_{\alpha}|^{2}}{2} dv_{\alpha} + \int f_{\beta} m_{\beta} \frac{|v_{\beta}|^{2}}{2} dv_{\beta} \end{bmatrix}$$

• Entropy growth:  $\partial_t \int f \ln f \, dv \ge 0$  when for example  $\partial_t f = S(f, f)$ .

Together these drive the distribution function towards the Maxwellian distribution:

$$M(x,t) = \frac{n(x,t)}{(2\pi T(x,t))^{d/2}} e^{-\frac{|v-u(x,t)|^2}{2T(x,t)}},$$

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# Motivation for reduced models

- The kinetic equations are extremely costly to simulate directly.
  - For a 3D3V problem, 100 degrees of freedom in each dimension means 10<sup>12</sup> grid points!
- Collisional diffusivity implies that hopefully not all those degrees of freedom are strictly needed in velocity space.
- Even much-reduced models can capture gross behavior, esp. flux of conserved quantities.

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- Even much-reduced models can capture gross behavior, esp. flux of conserved quantities.

#### Moment Methods

- Conservation laws posed only in physical space
- Includes classical fluid dynamics equations (Navier-Stokes, Euler)

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# Example: single species

Abbreviate the LHS by *Df*. We derive conservation of mass, momentum, and energy by integrating:

$$\begin{split} \langle Df &= S(f,f), 1 \rangle_{v} \implies \partial_{t} n + \nabla_{x} \cdot (nu) = 0, \\ \langle Df &= S(f,f), v \rangle_{v} \implies \\ \partial_{t}(nu) + \nabla_{x} \cdot (nT\overline{\overline{I}} + nu \otimes u + \mathbb{P}) = n(E + u \times B) \\ \langle Df &= S(f,f), |v|^{2}/2 \rangle_{v} \implies \\ \partial_{t} E + \nabla_{x} \cdot ((E + nT)u + \mathbb{P}u + q) = E \cdot u \end{split}$$

Fluid quantities are defined as<sup>1</sup>

$$n = \langle f \rangle_{v}, \quad u = \frac{1}{n} \langle f v \rangle_{v},$$
$$E = \langle f | v |^{2} / 2 \rangle_{v}, \quad T = \frac{1}{nd} \langle f | v - u |^{2} \rangle_{v}.$$

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<sup>1</sup>d is the dimension

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### Closure problem

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Unfortunately, the moment system is not closed.

#### Problem moments

• Anisotropic pressure tensor (2 dimensions):

$$\mathbb{P} = \left\langle f \begin{bmatrix} v_x^2 - |v|^2/2 & v_x v_y \\ v_y v_x & v_y^2 - |v|^2/2 \end{bmatrix} \right\rangle_{\mathbf{v}},$$

Heat flux:

$$\mathbf{q} = \langle f | \mathbf{v} - \mathbf{u} |^2 (\mathbf{v} - \mathbf{u}) \rangle_{\mathbf{v}}$$

#### ... lead to the moment closure problem

The task is to come up with expressions for  $\mathbb{P}$ , q (or their gradients), in terms of the lower order moments.

### Moment Closures

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- Accomplishing a moment closure requires simplification. Count the degrees of freedom!
- Two main "families" of theories.
  - Chapman-Enskog theory: explicitly perturbative about a Maxwellian distribution.
  - Ansatz for f: assume f belongs to some class of distributions or satisfies some variational principle.

#### Caution

The chosen moment closure should result in a hyperbolic system. (Eigenvalues of flux matrix)

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# Chapman-Enskog theory

Introduce a small parameter by assuming collisions are strong:  $\epsilon = Kn = \lambda/L$ .

$$\partial_t f_{\alpha} + \mathbf{v} \cdot \nabla_x f_{\alpha} + F \cdot \nabla_v f_{\alpha} = \frac{1}{\epsilon} \sum_{\beta} S(f_{\alpha}, f_{\beta})$$

Expand in powers of  $\epsilon$ , about  $f_{\alpha}^{0}$ :

$$f_{\alpha} = f_{\alpha}^{0} + \epsilon f_{\alpha}^{1} + \epsilon^{2} f_{\alpha}^{2} + \dots$$

Analyzing and retaining terms of a given order in  $\epsilon$  results in familiar fluid equations:

• Order 0: Euler equations (neutral), ideal two-fluid equations (plasma). "Everything is magically Maxwellian".

- Order 1: Navier-Stokes equations (neutral), Braginskii closure (plasma)
- Order 2: Burnett equations (unstable!)

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# Hermite polynomial expansion

Grad 1949 [6] used an expansion in Hermite polynomials,<sup>2</sup>

$$f = f^{0} \sum_{k=0}^{\infty} a^{k} \mathcal{H}^{k} \left( \frac{v - u}{v_{th}} \right)$$

- $\mathcal{H}^k$  are the tensorial rank-k Hermite polynomials,
- *a<sup>k</sup>* a corresponding tensor of coefficients.

• 
$$a^0 = 1, a^1_i = a^2_{ii} = 0$$

### Hermite polynomials?

The Hermite polynomials are orthogonal with respect to the Gaussian (Maxwellian) weight:

$$\int H^{i}(v)H^{j}(v)e^{-v^{2}/2} dv = \sqrt{2\pi}j!\delta_{ij}$$

#### Idea is to truncate the expansion.

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#### Benefits

- Capture the collisional asymptotic limit
- Reproduce conservation laws
- Focus computing resources on the quantities we care about: Grad's "physically meaningful" moments.
- Very natural to discretize using DG or finite volume methods.

#### Drawbacks

- Perturbative: moment convergence falls off quickly when far from a Maxwellian.
- Non-adaptive: no indication from the moments when your convergence is getting worse.

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# Moment methods recap

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# Brief history of dynamical low-rank methods

The dynamical low-rank method is a totally different direction for reduced plasma models.

- Proposed by Koch and Lubich for matrix-valued ODEs [7]
- Extended to higher-order tensors in various formats [8], [11]
- Overapproximation-insensitive integrators developed [10],
   [2]
- Applied to kinetic equations [5], [3]



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Example low-rank kinetic approximation

Consider an arbitrary equation of "kinetic type":

$$\partial_t f(x, v, t) = h$$

Look for approximate solutions of the form

$$\tilde{f} = \sum_{i,j=1}^{r} X_i(x,t) S_{ij}(t) V_j(v,t).$$

- The integer *r* is the rank
- Bases X and V are orthonormal:

$$\langle X_i, X_k \rangle_{\mathbf{x}} = \delta_{ik}, \quad \langle V_j, V_l \rangle_{\mathbf{v}} = \delta_{jl}$$

- Similar to the SVD, except that S is not necessarily diagonal.
  - $\implies$  the decomposition is not unique.

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# What makes it dynamical?

We impose a Galerkin condition on the time derivative of the system:

$$\langle h - \tilde{h}, \delta \tilde{f} 
angle = 0$$
 for all  $\delta \tilde{f} \in \mathcal{T}_{\tilde{f}}\mathcal{M}_r$ .

- $\mathcal{M}_r$  is the manifold of low-rank functions
- $\mathcal{T}_{\tilde{f}}\mathcal{M}_r$  is the tangent space to the manifold at  $\tilde{f}$ .
- $\tilde{h}$  is our low-rank time derivative.



Figure: The method takes steps along the low-rank manifold's tangent space.

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# Projector-splitting integrator

Equivalently,  $\tilde{h} = P(\tilde{f})h$  for an orthogonal projection  $P(\tilde{f})$ . It turns out that P has the form

$$\tilde{h} = P(\tilde{f})h = X_i \langle X_i, h \rangle_{x} - X_i \langle X_i V_j, h \rangle_{xv} V_j + \langle V_j, h \rangle_{v} V_j.$$

An operator splitting of the projector leads to a simple first order time integration scheme:

$$\begin{split} \tilde{f}' &= \tilde{f}(t_0) + \Delta t(V_j \langle V_j, h \rangle_v) \\ \tilde{f}'' &= \tilde{f}' - \Delta t(X_i \langle X_i V_j, h \rangle_{xv} V_j) \\ (t_1) &= \tilde{f}'' + \Delta t(\langle X_i, h \rangle_x X_i). \end{split}$$

The scheme is:

• Exact if f has rank r.

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 Robust to overapproximation: S can have vanishing singular values!

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Analysis

Time-complexity savings can be huge.

 $r \ll N_x, N_v \implies r^2(N_x + N_v) \ll N_x N_v$ 

<sup>3</sup>Einkemmer ([4]) presents a conservative low-rank scheme. 💿 🖉 🔊 🔍 🔿

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Analysis

Time-complexity savings can be huge.

 $r \ll N_x, N_v \implies r^2(N_x + N_v) \ll N_x N_v$ 

### Drawbacks

While an appealing technique, the dynamical low-rank method is not perfect.

- It fails to preserve mass, momentum, and energy.<sup>3</sup>
  - The approximation can "leak" conserved quantities into the truncated ranks.
- It fails to recover the asymptotic limit of the Maxwellian when collisions are strong.
  - This is because a Maxwellian with spatially varying parameters is not low-rank!

<sup>&</sup>lt;sup>3</sup>Einkemmer ([4]) presents a conservative low-rank scheme 💿 🖉 🔊 🤉

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# A simple collisional plasma model

Test case: high-field limit of the Vlasov-Ampere-Fokker-Planck system

$$\begin{cases} \partial_t f + \mathbf{v} \cdot \nabla_{\mathbf{x}} f + \frac{1}{\epsilon} E \cdot \nabla_{\mathbf{v}} f = \frac{1}{\epsilon} \nabla_{\mathbf{v}} \cdot (\mathbf{v} f + \nabla_{\mathbf{v}} f), \\ \partial_t E = -J \end{cases}$$

- The current density J is defined  $J = \langle vf \rangle_v$ .
- Physically, can describe high-frequency motion of electrons against a static ion background fluid.

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# Limiting distribution

Define the local "Maxwellian"

$$M(x,v,t) = e^{-\frac{|v-E(x,t)|^2}{2}}$$

Then our Vlasov-Fokker-Planck equation is

$$\partial_t f + \mathbf{v} \cdot \nabla_{\mathbf{x}} f = \frac{1}{\epsilon} \nabla_{\mathbf{v}} \cdot [M \nabla_{\mathbf{v}} (M^{-1} f)].$$

As  $\epsilon \rightarrow 0$ , f approaches

$$f=\frac{\rho(x)}{(2\pi)^{d/2}}M,$$

for the density  $\rho(x) = \langle f \rangle_v$ .

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# Limiting fluid system

What is the governing equation for  $\rho$  in the  $\epsilon \rightarrow 0$  limit? • Multiply by 1 and v and integrate:

$$\partial_t \rho + \nabla_x \cdot J = 0$$
  
$$\partial_t J + \nabla_x \cdot \langle v \otimes vf \rangle_v = \frac{1}{\epsilon} (\rho E - J).$$

Send 
$$\epsilon \to 0$$
:  
 $\rho E - J = 0 \implies \begin{cases} \partial_t \rho + \nabla_x \cdot (\rho E) = 0, \\ \partial_t E = -\rho E. \end{cases}$ 

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### The Asymptotic-preserving property

$$\begin{cases} C^{\text{kinetic}}: \quad \partial_t f + v \cdot \nabla_x f = \frac{1}{\epsilon} \nabla_v \cdot [M \nabla_v (M^{-1} f)], \\ C^{\text{fluid}}: \quad f = \frac{\rho}{(2\pi)^{d/2}} M, \text{ where } \partial_t \rho + \nabla_x \cdot (\rho E) = 0 \end{cases}$$

Both coupled with Ampere's equation. We're looking for a kinetic discretization  $D^{\text{kinetic}}$  that makes this diagram commute:



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**Goal:** the bottom path should hold for fixed r.

Low-rank algorithm

Key idea: evolve low-rank representation of  $g = M^{-1}f$ :

$$\epsilon o 0 \quad \Longrightarrow \quad g o rac{
ho(x)}{(2\pi)^{d/2}} \; ({
m rank} \; 1)$$

Recall the low-rank ansatz is

$$g(x,v,t) = \sum_{i,j=1}^{r} X_i(x,t)S_{ij}(t)V_j(v,t).$$

Plug f = gM into the Vlasov-Fokker-Planck equation:

$$h \triangleq \partial_t g = -v \cdot \nabla_x g - \frac{1}{M} (\partial_t M + v \cdot \nabla_x M) g \\ + \frac{1}{\epsilon} [(\nabla_v - v + E) \cdot \nabla_v g].$$

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### Projector-splitting integration

Introduce auxiliary bases K and L:

$$X_j(x,t) = \sum_i X_i(x,t)S_{ij}(t), \quad L_i(v,t) = \sum_j S_{ij}(t)V_j(v,t).$$

**1** Advance 
$$K_j^n$$
 to  $K_j^{n+1}$  via  $\partial_t K_j = \langle V_j, h \rangle_v$ .

2 Perform a QR decomposition of  $K_i^{n+1}$  to obtain  $X^{n+1}, S'$ .

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- 3 Advance S' to S" via  $\partial_t S_{ij} = -\langle X_i V_j, h \rangle_{xv}$ .
- 4 Advance  $L_i^n$  to  $L_i^{n+1}$  via  $\partial_t L_i = \langle X_i, h \rangle_x$ .
- **5** Perform a QR decomposition of  $L_i^{n+1}$  to obtain  $V^{n+1}, S^{n+1}$ .

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### Spatial and temporal discretization

- First-order finite difference discretization in X and V
- First-order timestepping with implicit treatment of stiff collision terms
- Caution is required with collision term in "backwards" S step. Instability in time-reversed diffusion equation can be partly tamed with Forward Euler cancellation. E.g.  $dy/dt = \lambda t$ :

$$egin{aligned} y_1 &= (1-\lambda\Delta t)^{-1}y^n \ y_2 &= (1+\lambda(-\Delta t))y_1 = y^n \ y^{n+1} &= (1-\lambda\Delta t)^{-1}y_2 = (1-\lambda\Delta t)^{-1}y^n. \end{aligned}$$

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For very small  $\epsilon$ , we capture the high field limit using only a small rank.





Figure: First-order convergence to the fluid limit in  $\Delta x$ 

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# Singular value evolution

The singular values of the solution relax in one time step to order  $\boldsymbol{\epsilon}$ 

Magnitude of singular values

10° 10° 10-5-10-5-10-10-10-10-- σ\_3 — σ\_4 10-15-10-15σ5 0.000 0.001 0.000 0.025 0.050 0.075 0.100

Figure: Singular values of solution beginning in local equilibrium

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### Singular value evolution

The singular values of the solution relax in one time step to order  $\boldsymbol{\epsilon}$ 



Figure: Singular values of solution beginning far from local equilibrium

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# Ratio of singular values

Ratio of first two singular values



Figure: The ratio of the two largest singular values relaxes to order  $\epsilon$  in one timestep.

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### Larger $\epsilon$ requires higher rank



Figure: Higher rank is required to resolve less collisionless dynamics.

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 ${\sf Conclusion}$ 

- Dynamical low-rank methods can capture the collisional fluid limit of kinetic equations.
- Arguably more flexible than moment methods, but less structure-preserving.

#### Next steps

- Application to magnetized, multispecies plasmas.
- Attempt to enforce conservation following Einkemmer [4].
- Introduce adaptivity by increasing rank as necessary.

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